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Physical interpretation of the Weert superpotential

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Abstract. Weert [1] obtained a potential for the bounded part of the Maxwell Tensor which is associated to the Liénard-Wiechert field. We show that this potential can be interpreted as an intrinsic Angular Momentum Density for the corresponding Electromagnetic Field.

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1. Introduction

In this paper we deal with electromagnetic field which is produced by a point particle charge in the Minkowski space. This charge gives the Liénard-Wiechert field (LW) from which the Maxwell tensor T_{jc} can be obtained. The last can be separated in two parts: T_{jc}^B the bounded one, and T_{jc}^R the radiative part (in the sense of Teitelboim [2]).

Sec. 2 is devoted to a brief exposition of the LW field. Weert [1] constructed a potential K_{jbc}^B for T_{jc}^B , this will serve us to propose, in Sec. 3, a physical interpretation of this potential using the Bhabha [4]-Synge [5] region; moreover, if we decompose K_{jbc}^B in two parts which satisfy the symmetry of the Lanczos Spintensor [6], we obtain the rupture of T_{jc}^B proposed by López [9], which is very important in the study of the angular momentum of the LW field. In Sec. 4 we construct a non-local Superpotential (depending of the past history of the charge) for the radiative part; and we show the terms of T_{jc}^R which do not participate in the flux of energy and momentum through the Bhabha-Synge tube.

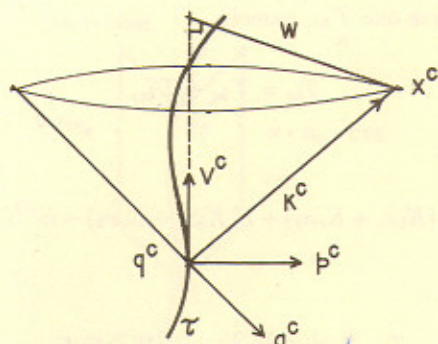


FIGURE 1. Kinematics of the world-line.

2. Point charge in arbitrary motion

A charged particle moving in an arbitrary motion in the Minkowskian space, produced the electromagnetic LW field with the 4-potential and Faraday tensor given by

$$A^c = qw^{-1}v^c, \quad F_{bc} = qw^{-2}(U_bK_c - U_cK_b), \quad (1.a)$$

and the Maxwell Tensor

$$T_{bc} = q^2w^{-4} \left[K_bU_c + K_cU_b + (a^2 - B^2)K_bK_c + \frac{1}{2}g_{bc} \right], \quad (1.b)$$

where (Fig. 1):

$$(x^c) = (x, y, z, t), \quad (g_{bc}) = \text{Diag}(1, 1, 1, -1)$$

τ is the proper timer, $q^c(T)$ will be the retarded point, v^c is the 4-velocity, a^c the 4-acceleration

$$K^c = x^c - q^c \quad (1.c)$$

$w = -K^c v_c$ which is the so called retarded distance

$$W = -K^c a_c, \quad a^2 = a^c a_c$$

$B = w^{-1}(1 - W)$ known as the Plebański [10] invariant

$$U_c = Bv_c + a_c, \quad p^c = w^{-1}K^c - v^c.$$

Teitelboim [2] proved that Eq. (1.b) can be written as the sum of a bounded part T_{Bbc} and a radiative one T_{Rbc} , namely,

$$T_{bc} = T_{Bbc} + T_{Rbc}, \quad (2.a)$$

such that

$$T_{Bbc} = q^2 w^{-4} \left[\frac{1}{2} g_{bc} + (K_b a_c + K_c a_b) + B(K_b v_c + K_c v_b) - w^{-2} (1 - 2W) K_b K_c \right], \quad (2.b)$$

and

$$T_{Rbc} = q^2 w^{-2} (\alpha^2 - w^{-2} W^2) K_b K_c. \quad (2.c)$$

Each tensor possesses the following differential properties

$$T_{Bb,c}{}^c = 0, \quad (2.d)$$

$$T_{Rb,c}{}^c = 0, \quad (2.e)$$

respectively, when it is valued out of the world-line.

Weert [1] proved that Eq. (2.d) comes out from the existence of the superpotential

$$K_{Bjbc} = -q^2/4w^{-4} \left[w^{-1} (3 - 4W) (v_j \times K_b) K_c + 4(a_j \times K_b) K_c + g_{cj} K_b - g_{cb} K_j \right], \quad (3.a)$$

where we have used the Lowry [3] notation

$$A_j \times B_c = A_j B_c - A_c B_j, \quad (3.b)$$

such that

$$T_{Bbc} = K_{Bb}{}^j{}_{c,j}. \quad (3.c)$$

From here, (2.d) follows immediately.

In the next section we will study (3.a) in order to give a physical interpretation of this superpotential. In Sec. 4 we will do a brief analysis of Eqs. (2.c-e).

3. Weert superpotential

Our point charge in arbitrary motion gives an electromagnetic field which possesses an intrinsic angular momentum (IAM). Here we will prove that K_{Bjbc} behaves like a

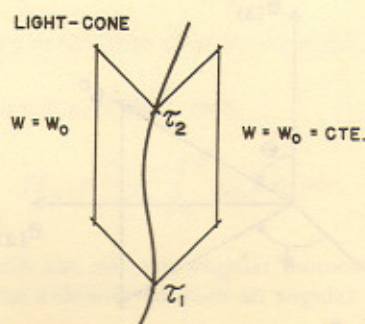


FIGURE 2. Bhabha-Synge tube.

density for IAM when the corresponding fluxes are calculated through a Bhabha [4]-Synge [5] tube.

The superpotential (3.a) has the following properties

$$\begin{aligned}
 K_B{}^{jbc} &= -K_B{}^{bjc} && \text{antisymmetry,} \\
 K_B{}^j{}_c{}^c &= 0 && \text{null-trace,} \\
 K_B{}^{jbc} + K_B{}^{bcj} + K_B{}^{cjb} &= 0 && \text{cyclic,} \\
 K_B{}^{jbc}{}_{,c} &= 0 && \text{null divergence.}
 \end{aligned} \tag{4}$$

These properties agree with those of the Lanczos [6] Generator K_{jbc} for the Weyl tensor of the space-time. Lanczos calculated K_{jbc} for weak gravitation fields and in his analysis, the Dirac equation, for spin $\frac{1}{2}$, appeared. For this reason he called the potential K_{jbc} spin-tensor. In our case, $K_B{}^{jbc}$ will be associated with the IAM of the LW field.

Consider the Bhabha-Synge tube (Fig. 2), which is composed by the light cones with the tops in $\tau = \tau_1$ and $\tau = \tau_2$, and a surface of constant retarded distance. First, let us calculate the $K_B{}^{jbc}$ flux through a light cone; the expression is given by Synge [5]

$$\tilde{M}_{jb} = \int_{\tau=\text{const}} K_B{}^{jbc} d\sigma^c = - \int_0^{w_0} w dw \int_{\tau=\text{const}} d\Omega K_B{}^{jbc} K^c, \tag{5.a}$$

where $d\Omega$ is the element of solid angle.

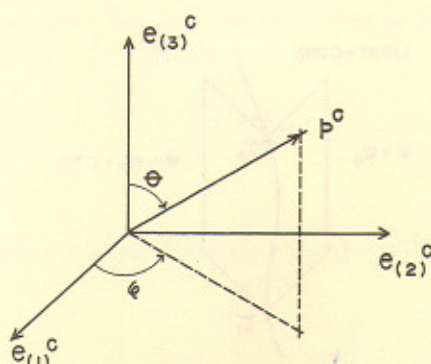


FIGURE 3. Fermi tetrad.

The unitary space-like vector p^c has been defined in Fig. 1

$$p^c = w^{-1} k^c - v^c, \quad p^c p_c = 1, \quad p^c v_c = 0. \quad (5.b)$$

Now for any event in the line-universe, p^c can be written in terms of a Fermi tetrad $e_{(\gamma)c}$ with $\gamma = 1, 2, 3$ (Fig. 3).

$$\begin{aligned} \frac{de_{(\gamma)}^c}{d\tau} &= a_{(\sigma)} v^c = (a^r e_{(\sigma)r}) v^c, \\ p^c &= \sin \theta \cos \phi e_{(1)}^c + \sin \theta \sin \phi e_{(2)}^c + \cos \theta e_{(3)}^c \\ &= P_{(\sigma)} e_{(\sigma)}^c = (p^r e_{(\sigma)r}) e_{(\sigma)}^c \end{aligned} \quad (5.c)$$

and $d\Omega = \sin \theta d\theta d\phi$.

From (3.a), it is clear that

$$K_{jbc} K^c = 0. \quad (5.d)$$

Therefore (5.a) implies $\tilde{M}_{jb} = 0$, that is, K_{jbc} flux vanishes through a light cone.

Now we will calculate the K_{jbc} flux through the 3-space $w = w_0 = \text{const}$; the expression is given in Synge [5]

$$M_{jb} = \int_{w=w_0} K_{jbc} d\sigma^c = w^2 \int_{\tau_1}^{\tau_2} d\tau \int K_{jb}{}^c w_{,c} d\Omega, \quad (6.a)$$

where

$$w_{,c} = \text{Gradient of } w = -v_c + BK_c \quad (6.b)$$

Therefore, Ecs. (3.a, 5.c, 6.a, b) imply that

$$M_{I\ jb} = \frac{8\pi}{3} q^2 \int_{\tau_1}^{\tau_2} (v_j \times a_b) d\tau. \quad (6.c)$$

This last result agrees with the intrinsic angular momentum of the LW Field [7]. Therefore, $K_{B\ jbc}$ behaves as a density for such an angular momentum.

The superpotential (3.a) accepts the following rupture

$$K_{B\ jbc} = \tilde{K}_{B\ jbc} + \bar{K}_{B\ jbc}, \quad (7.a)$$

where

$$\tilde{K}_{B\ jbc} = q^2 w^{-4} [(-a_j + w^{-1} W v_j) \times K_b] K_c \quad (7.b)$$

and

$$\bar{K}_{B\ jbc} = -w^{-4} [g_{cj} K_b - g_{cb} K_j + 3w^{-1} (v_j \times K_b) K_c]. \quad (7.c)$$

The potentials in Eqs. (7.b, c) satisfy all the properties in (4), as the Lanczos Spin-tensor does.

It is simple to prove that

$$\int_{\tau=\text{const}} \bar{K}_{B\ jbc} d\sigma^c = \int_{w=\text{const}} \bar{K}_{B\ jbc} d\sigma^c = 0, \quad (8.a)$$

and

$$M_{I\ jb} = \int_{w=\text{const}} \tilde{K}_{B\ jbc} d\sigma^c \quad \text{and} \quad \int_{\tau=\text{const}} \tilde{K}_{B\ jbc} d\sigma^c = 0. \quad (8.b)$$

That is, $\bar{K}_{B\ jbc}$ doesn't contribute to the IAM of the electromagnetic field; this means that $\tilde{K}_{B\ jbc}$ is the active part of part of $K_{B\ jbc}$.

The result can be obtained by using Stokes Theorem and the Rowe [8] identity

$$\bar{K}_{B\ jbc} = \left(\frac{q^2}{4} w^{-4} D_{jbc} r \right)_{,r}, \quad (8.c)$$

where D_{jbc} is a tensor used by Sygne [5] in another context

$$D_{sarb} = (g_{sr}K_b - g_{sb}K_r)K_a + (g_{ab}K_r - g_{ar}K_b)K_s. \quad (8.d)$$

Finally using (7.a) in (3.c), the following decomposition is obtained

$$T_{Bbc} = \tilde{T}_{Bbc} + \bar{T}_{Bbc}, \quad (9)$$

with $\tilde{T}_{Bbc} = \tilde{K}_{Bbc,j}^j$ and $\bar{T}_{Bbc} = \bar{K}_{Bbc,j}^j$.

Eq. (9) is important at the moment we relate it to the angular momentum radiated by the charge [9].

4. On a rupture for T_{ab} R

Here we show how the radiative part of T_{ab} can be written as the sum of two terms; one of them doesn't participate in the energy and momentum flux through the Bhabha-Syngé tube. Moreover, we will give a potential for T_{Bbc} .

The expression (2.c) can be written in the form

$$T_{Rbc} = T_{bc} + \tilde{T}_{bc}, \quad (10.a)$$

where

$$T_{bc} = q^2 w^{-4} (a^2 - 3w^{-2}W^2) K_b K_c, \quad T_{b,c}^c = 0 \quad (10.b)$$

$$T_{Rbc} = 2q^2 w^{-6} W^2 K_b K_c, \quad T_{Rb,c}^c = 0 \quad (10.c)$$

It is simple to demonstrate that

$$\int_{\substack{\tau = \text{const} \\ \text{or} \\ w = \text{const}}} T_{bc} d\sigma^c = 0, \quad (11.a)$$

that is, (10.b) doesn't contribute to the energy and momentum flux through the Bhabha-Syngé tube. In a similar way

$$\int_{\substack{w = \text{const} \\ \text{or} \\ \tau = \text{const}}} (x^j T_{bc}^j - x^b T_{bc}^j) d\sigma_c = 0. \quad (11.b)$$

Hence T_{bc} doesn't participate either in the momentum fluxes for such tube. Due to Ecs. (11.a, b) we say that the tensors in (10.b) represent the inactive part of $T_{\mathbb{R}}^{ab}$ with respect to the Bhabha-Syngé region.

The conservation law is immediately deduced from the existence of the superpotential:

$$K_{i jbc} = -\frac{q^2}{4} w^{-2} \left[w^{-2} W^2 (g_{cj} K_b - g_{cb} K_j) + w^{-1} W (v_j \times K_b) (a_c - 3w^{-2} W K_c) \right. \\ \left. + (a_j \times K_b) (4w^{-2} W K_c - a_c) \right], \quad (12.a)$$

such that

$$T_{bc} = K_{i b c, j}^j. \quad (12.b)$$

Moreover, the identity $T_{\mathbb{R}}^{bc}{}_{,c} = 0$ is a consequence of

$$\tilde{T}_{bc} = \tilde{K}_{i b c, j}^j, \quad (13.a)$$

where

$$\tilde{K}_{i b j c} = -2q F_{bj} p(\sigma) p(\gamma) \left[\int_0^\tau a(\sigma) a(\gamma) v_c d\tau + p(\beta) \int_0^\tau a(\sigma) a(\gamma) e_{(\beta)c} d\tau \right] \quad (13.b)$$

in equation (13.b), the sum over $\sigma, \gamma, \beta = 1, 2, 3$ has to be done.

To verify Eq. (13.a), we have to use the following relations

$$\begin{aligned} \tau_{,j} &= -w^{-1} K_j, && \text{retarded derivative} \\ F_b^j{}_{,j} &= 0, && \text{Maxwell equations} \\ F_b^j{}_{,j} &= qw^{-2} K_b, && \text{null eigenvector} \\ F_b^j p(\sigma)_{,j} &= 0, && \text{Fermi tetrad} \\ W &= -wp(\sigma)a(\sigma), \\ k^c &= w(v^c + p^c). \end{aligned} \quad (13.c)$$

The existence of integrals in equation (13.b) shows the non-local character of $\tilde{K}_{\mathbb{R}}^{bjc}$: that means that it depends on the past history of the charge. Therefore

Eqs. (2.a, 3.c, 10.a, 12.b, 13.a) imply

$$T_{bc} = \left(K_{B \ b \ c}^j + K_{i \ b \ c}^j + \tilde{K}_{R \ b \ c}^j \right)_j. \quad (14)$$

Hence the Maxwell tensor associated to the LW Field is an *exact divergence*.

References

1. Ch.G. Van Weert, *Phys. Rev.* **D9** (1974) 339.
2. C. Teitelboim, *Phys. Rev.* **D1** (1970) 1572.
3. E.S. Lowry, *Phys. Rev.* **117** (1960) 616.
4. H.J. Bhabha, *Proc. Roy. Soc. London* **A172** (1939) 381.
5. J.L. Synge, *Ann. Math. Pura Appl.* **84** (1970) 33.
6. C. Lanczos, *Rev. Mod. Phys.* **34** (1962) 379.
7. C.A. López and D. Villarroel, *Phys. Rev.* **D11** (1975) 2724.
8. E.G.P. Rowe, *Phys. Rev.* **D18** (1978) 3639.
9. C.A. López, *Phys. Rev.* **D17** (1978) 2004.
10. J. Plebański, The structure of the field of point charge. Reporte Interno. Departamento de Física CIEA-IPN (1972).

Resumen. Weert [1] obtuvo un potencial para la parte acotada del tensor de Maxwell asociado al campo de Liénard-Wiechert. Aquí mostramos que dicho potencial puede interpretarse como una densidad de momento angular intrínseco del correspondiente campo electromagnético.